| 1  | TERRESTRIAL RECORD OF CYCLIC EARLY EOCENE WARM-HUMID EVENTS IN CLAY MINERAL   |
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| 2  | ASSEMBLAGES FROM THE SALTA BASIN, NORTHWESTERN ARGENTINA  |
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## 24 ABSTRACT

25 The Eocene continental sequence investigated in this study belongs to the Salta Group; it was 26 deposited in an intracontinental rift, the Salta Basin (NW Argentina) that evolved from the Lower 27 Cretaceous to middle Paleogene. The Salta Group contains the Maíz Gordo and Lumbrera Fms, spans 28 the Paleocene-early Eocene, and shows excellent exposures in the region of the Valles Calchaquíes. 29 This research is focused on the continental facies of the Lumbrera Fm, which correspond to the early 30 Eocene. We studied the mineralogy of the fine-grained beds of the Lumbrera Fm in five locations 31 (Valle Encantado, Tonco, Tin Tin, Luracatao, and Obelisco) by X-ray diffraction and scanning electron 32 microscopy in order to examine the vertical variations in clay mineralogy and their relations with global paleoclimatic changes registered during the Eocene. The microtexture of the authigenic 33 34 smectite-type clays (Sm to I/Sm R0) suggests that they mainly originated by crystallization from 35 glassy volcanic materials. The high reactivity of the glass precludes the use of smectite-type-clay formation as an indicator of specific paleoclimatic conditions in the studied sediments. In contrast, 36 37 the formation of kaolinite in sections with very low smectite proportions and a strong degree of 38 weathering in crystalline silicates reflects intense weathering conditions and is a useful paleoclimatic 39 proxy in terrestrial sediments. Significant variations in kaolinite/muscovite ratios at the base and in an intermediate bed in the Lumbrera Fm at Valle Encantado suggest the presence of cyclic 40 41 hyperthermals through the Ypresian stage that may be tentatively correlated with the Eocene 42 Thermal Maxima 2 and 3, which are the largest events that have been identified at a global scale in 43 early Eocene marine sediments.

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45 **KEYWORDS:** Paleoclimate reconstruction, continental basins, kaolinite, EECO

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### 48 1. INTRODUCTION

From the late Paleocene (58 Ma) to the early Eocene the Earth's surface underwent a short 49 50 period of warming associated with elevated levels of atmospheric CO<sub>2</sub>; according to various lines of 51 evidence, this was the warmest period in the Cenozoic Era (Kennett and Stott, 1991; Zachos et al., 52 1993, 2001; Schmitz and Pujalte, 2007; Sluijs et al., 2007; McInerney and Wing, 2011). This event, 53 known as the Paleocene-Eocene thermal maximum (PETM or ETM-1), occurred ~56 million years ago 54 in a short period of time that spanned ~170,000 years (Röhl et al., 2007; Westerhold, 2008; Charles 55 et al., 2011; Zeebe et al., 2016). A few million years later there was another long-term warming trend 56 forced by high concentrations of atmospheric CO<sub>2</sub>, which comprises a period called the Early Eocene Climatic Optimum (EECO; ~53-49 Ma; Zachos et al., 2001, 2008; Kirtland Turner and Ridgwell, 2013; 57 58 Lauretano et al., 2015, 2018; Westerhold et al., 2018; Crouch et al., 2020). This long-term trend 59 through the Ypresian stage was punctuated by recurrent transient warming events, associated with 60 carbon isotope excursions (CIEs), called "hyperthermals" (Thomas, 1998; Thomas et al., 2000; Zachos et al., 2010). The chronology of the different hyperthermals, their relationship with eccentricity 61 cycles, as well as the evolution of surface and sea-bottom temperatures, have been established 62 63 mainly from the study of deep-sea cores (Cramer et al., 2009; Zachos et al., 2010; Laurentano et al., 2015, 2018; Thomas et al., 2018; Westerhold, et al., 2018; Crouch et al., 2020). The hyperthermal 64 events of the Paleocene-early Eocene were initially recognized in several deep-sea cores from the 65 66 Atlantic Ocean (e.g., Cramer et al., 2003; Lauretano et al., 2015; Littler et al., 2014, 2010; Zachos et 67 al., 2005, 2010), but were subsequently also identified in deep-sea cores from the Pacific Ocean as 68 well as in marine sedimentary successions outcropping in various localities (Nicolo et al., 2007; 69 Galeotti et al., 2017; Westerhold et al., 2018). Initially two warming events, ETM-2 and 3, were 70 identified in the early Eocene (Zachos et al., 2010; Röhl et al., 2005), but the high-resolution analysis of deep-sea cores identified several prominent CIEs occurring between 55 and 52 Ma, which were 71 72 labeled from E to L by Cramer et al. (2003) and have been correlated with short eccentricity cycle maxima. Recently, excursions H to L, as well as 20 additional smaller CIEs, have been documented for the Ypresian to early Lutetian (56–45 Ma) in stable isotope records from sea cores from the Demerara Rise in the equatorial Atlantic (Kirtland Turner and Ridgwell, 2013; Sexton et al., 2011; Laurentano et al., 2016). The paleoclimatic changes taking place during this period have been extensively studied in the last few decades since they are considered analogous to current global greenhouse conditions.

79 At the same time, the study of the climatic conditions prevailing during the early Eocene in 80 terrestrial realms is crucial to understanding how the current global greenhouse climate could affect 81 temperatures and hydrological cycles in inner continental areas. In fact, studies in different basins 82 indicate that early Eocene hyperthermal events were associated with marked changes in hydrological cycles, which in turn modified surface processes such as the rate of erosion and chemical weathering 83 84 (Clechenko et al., 2007; Schmitz and Pujalte, 2007; Bataille et al., 2019). However, our knowledge of 85 continental climates during the Paleocene-early Eocene remains limited. There is a consensus that the global climate was warmer than at present during the Mesozoic and Early Cenozoic, but 86 87 disagreements remain with respect to winter temperatures and latitudinal gradients, and also the 88 consequences of warming events on hydrological cycles (Greenwood and Wing, 1995). These studies 89 have applied different approaches to understanding the environmental conditions that prevailed 90 during the early Eocene, including the use of carbon and oxygen isotope data, paleontological data 91 (i.e., foliar physiognomy and floristic composition), the characterization of paleosols, the study of 92 clay assemblages, and sedimentological evidence (Greenwood and Wing, 1995, 2010; Bataille et al., 93 2016, 2018; Andrews, et al., 2017; Do Campo et al., 2018; Kelson et al., 2018). However, most of the 94 studies that have focused on terrestrial sections have been carried out in the Northern Hemisphere 95 (Koch et al., 1995, 2003; Kraus, 1997; Domingo et al., 2009; Hyland and Sheldon, 2013; Bataille et al., 96 2016; Kelson et al., 2018; Song et al., 2018; Willard et al., 2019), and only a few in South America 97 (Krause et al., 2010; Hyland and Sheldon, 2015, 2017; Andrews et al., 2017).

98 The Santa Barbara Subgroup (Salta Group) represents an interesting case study to ascertain the changes in paleoclimate that took place in the Paleocene-early Eocene, since its upper two units, the 99 100 Maíz Gordo and Lumbrera Fms, correspond to this period and provide excellent exposures in the 101 region of Valles Calchaquíes (Fig. 1). The paleoclimatic changes that occurred during the deposition 102 of the Maíz Gordo Fm have been the subject of several studies (del Papa, 1999; Do Campo et al., 103 2007, 2018; Andrews et al., 2017). By contrast, specific studies of the Lumbrera Fm have not yet been 104 carried out, and the few existing studies have addressed this unit only briefly (del Papa and Salfity, 105 1999; Andrews et al., 2017; Do Campo et al., 2018). The aim of this study is thus to undertake a 106 characterization of the clay mineral assemblages together with a sedimentological analysis of the 107 Lumbrera Fm in several sites corresponding to the western part of the Salta Basin (northwestern 108 Argentina). To this end, we integrate X-ray diffraction (XRD) analysis with electron microscopy images 109 in order to understand the textural relations and origin of clays in the Lumbrera Fm and, what is 110 more, to correlate horizontal and vertical changes in clay mineral assemblages with global 111 paleoclimate changes taking place during the early Eocene.

112 The formation of clay minerals at the earth's surface proceeds by weathering and authigenic 113 reactions that depend on the complex interplay of a number of variables, including climate, source-114 area lithology, continental morphology, and the depositional environment (Chamley, 1989). 115 However, if these variables remain constant for enough time the genesis of clay minerals will be 116 mainly controlled by weathering intensity, which is in turn controlled by climatic conditions. When 117 this occurs, clay minerals reach a state close to equilibrium with their environment and are thus 118 representative of the climatic conditions prevailing during their formation in soil profiles (Thiry, 119 2000). For these reasons, clay mineral assemblages of continental sequences have been successfully 120 used to infer paleoclimatic and paleoenvironmental conditions in Mesozoic and Cenozoic strata (e.g., 121 Ruffell et al., 2002; Raucskik and Varga, 2008; Bauluz et al., 2014; Do Campo et al., 2018 and 122 references therein).

123 Under dry and cold climates physical weathering prevails, generating clay mineral assemblages 124 dominated by illite and/or chlorite. Under warm climate conditions with alternating humid and dry 125 seasons, chemical weathering is moderate giving rise to smectite formation (Buurman et al., 1988; 126 Güven, 1988; Chamley, 1989; Murakami et al., 1996). By contrast, humid subtropical to tropical 127 climates are related with highly hydrolytic conditions that in turn promote kaolinite formation in 128 sediments and soil profiles (Chamley, 1989; Righi and Meunier, 1995). Kaolinite usually forms 129 through the dissolution of aluminosilicates such as feldspars or micas in the presence of water, and 130 these reactions are enhanced by high temperatures and low pH, because under acid conditions silica 131 will pass into the aqueous solution more than alumina (Velde, 1992). In view of the contrasting 132 climatic conditions that favor illite and kaolinite genesis in weathering profiles, Chamley (1989) proposed the kaolinite/mica ratio (Kln/Ms), calculated from their relative abundances, as a 133 134 paleoclimate proxy.

135 However, syn-sedimentary volcanic events could also leave a significant mark in clay mineral 136 assemblages, as the highly labile volcaniclastic material that ends up in a basin can easily be 137 transformed into smectite during early diagenesis (Cuadros et al., 1999; Do Campo et al., 2010). In 138 the case of the Salta Basin, which developed to the east of an active volcanic arc, volcanic input should be considered as a possible source material. Accordingly, standard XRD analysis was 139 140 complemented with a detailed study of representative samples by scanning electron microscopy 141 employing backscattered electron images (BSE) and secondary electron images (SE), together with 142 established methodologies of sedimentary facies analysis. In a previous study we have demonstrated 143 that textural and morphological analysis of fine-grained sediments by SEM is essential to discriminate 144 between clay minerals formed through physical and chemical weathering in soil profiles and those 145 formed by early diagenesis of volcaniclastic material (Do Campo et al., 2010).

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#### 149 2. METHODS AND SAMPLES

Stratigraphic sections were measured and described bed by bed in five locations: Valle Encantado, Tin Tin, and Tonco, which are situated in the Parque Nacional Los Cardones; Obelisco, which is located near Cafayate; and Luracatao–Represa, which is situated to the west in Cumbres de Luracatao (Fig. 1). The aim was to study the changes in clay mineral associations in fluvial environments. Each site was sampled taking into account the characteristics of the sedimentary facies of the Lumbrera Formation. It is worth mentioning that the Tin Tin and Obelisco sections coincide with those reported in Andrews et al. (2017).

157 The mineralogical composition of 87 whole rocks and clay sub-samples from the Lumbrera Fm 158 (32 from Valle Encantado, 18 from Tonco, 18 from Tin Tin, 12 from Luracatao-La Represa, and seven 159 from Obelisco) was analyzed by X-ray diffraction (XRD); the distribution of the samples along the 160 stratigraphic logs is shown in Figures 2 and 3. XRD was performed with a PANalytical X'Pert Pro 161 diffractometer (CuKa radiation, 45kV, 40mA) equipped with an X'Celerator solid-state linear detector 162 (Department of Mineralogy and Petrology, University of Granada). Clay sub-samples (<2 µm) were 163 prepared in accordance with the guidelines of Moore and Reynolds (1997), and the mineral intensity 164 factors (MIF) of the same authors were employed to estimate the relative abundances of illite-mica, 165 kaolinite, chlorite, smectite, and I/Sm mixed-layer clays. The ratio of kaolinite and illite-mica (Kln/Ms) 166 abundances was then calculated.

We performed standard petrographic analyses to characterize the lithology and textures of the samples corresponding to the five study sections. Afterwards, sixteen samples characterized by their high kaolinite or smectite content were chosen for detailed study using field scanning electron microscopy (FESEM), employing a Carl Zeiss FESEM to obtain textural and chemical information on the clay minerals at micro and nanoscale. In the case of the backscattered electron study (BSE) and the energy-dispersive X-ray (EDS) analysis, polished thin sections were employed. The accelerating voltage used was 15Kv with a beam current of 1 nA. Atomic concentration ratios obtained by EDS

were converted into formulae according to stoichiometry. Accordingly, the structural formulae of
 dioctahedral smectite were calculated on the basis of 22 negative charges (O<sub>10</sub>(OH)<sub>2</sub>).

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## 177 **3. GEOLOGICAL SETTING**

During Cretaceous-Eocene times, an intracontinental rift basin – the Salta Basin – developed in northwestern Argentina (Salfity and Marquillas, 1994; Viramonte et al., 1999). The successions that were deposited make up the Salta Group, which consists, from base to top, of the Pirgua, Balbuena and Santa Bárbara Subgroups. The Santa Bárbara Subgroup is in turn constituted by the Mealla, Maíz Gordo, and Lumbrera formations, which comprise fluvial environments in western proximal areas and lakes in the center of the basin (for details, see del Papa and Salfity, 1999).

The uppermost Lumbrera Fm, the focus of this study, is in paraconformable contact with the underlying Maíz Gordo Fm and is unconformably covered by the overlying Quebrada de los Colorados Fm of the foreland Payogastilla Group (del Papa, 2006).

187 Age constraints on the units integrating the Santa Bárbara Subgroup are mainly based on 188 biostratigraphic and palynostratigraphic studies. Land-mammal associations and palynostratigraphy 189 indicate that the Mealla and Maíz Gordo formations date to the Selandian and Thanetian respectively, the final two stages of the Paleocene Epoch, whereas the Lumbrera Fm has been 190 191 assigned to the Eocene (Pascual et al., 1981; Quattrocchio et al., 2005; del Papa et al., 2010). 192 Moreover, in the upper ~30 m of the Maíz Gordo Fm, comprising a thick prominent paleosol section, 193 Andrews et al. (2017) identified several prominent CIEs from -22‰ to ~-28‰, which they correlated 194 with the PETM, thus indicating that this interval spans the Paleocene-Eocene transition.

Additionally, the mammal fossils (isotemnids) recovered from basal levels of the Quebrada de Los Colorados Fm in Luracatao Valley, as well as a maximum depositional age of ~37Ma inferred from U-Pb ages from detrital zircons at Angastaco, indicate an Eocene age for this unit, probably middlelate Eocene (Payrola et al., 2009; Carrapa et al., 2012), thus constraining the age of the Lumbrera Fm

to the early Eocene. By contrast, a recent study carried out on nearby coeval sections applied
paleomagnetic results to establish the bottom of the Maíz Gordo Fm in the Ypresian (Hyland et al.,
2015, 2017), the first stage of the Eocene Epoch. In reply to a comment from Hyland and Sheldon
(2017), however, White et al. (2017) have questioned the Hyland et al. (2017) approach to
constraining their magnetostratigraphic record and therefore the age model they proposed.

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# 205 4. SEDIMENTARY FACIES OF THE LUMBRERA FORMATION

The present study focuses on the fluvial environment of the Lumbrera Fm, mainly consideringthe fine-grained lithologies of the floodplain setting (Figs. 2, 3).

In the study area, the thickness of the Lumbrera Fm ranges from a maximum of 237 m in the Valle Encantado section to a minimum of 60 m in the Luracatao-La Represa section, with intermediate values in the Tin Tin, Obelisco, and Tonco sites (187, 169 and 121 m, respectively) (Figs. 2, 3). The Obelisco section is not shown because only control samples were taken in this locality.

The Lumbrera Fm consists of well-developed stacked channel facies interbedded with finegrained floodplain facies (Fig. 4A). The sedimentary facies display lateral changes from west to east, with conglomerates dominating in the western sections Luracatao-La Represa (Luracatao hereafter), Tin Tin, and Tonco-La Escalera (Tonco hereafter) (Figs. 2B and 3), and sandy lithologies prevailing in the Valle Encantado section (Fig. 2). The main facies and facies associations are summarized in Table 1.

In the western area (the Luracatao and Tin Tin sections), channel-fill facies associations (FA1) are characterized by shallow lenticular geometries 2-3 m thick. They are composed of medium to fine, white to pinky conglomerates with normal gradation and trough cross-stratification (Fig. 4C). The tops of these beds display bioturbation and red to purple mottles (Fig. 4B). The fine-grained floodplain deposits (FA3) associated with the conglomerates consist of tabular, red to brown sandy siltstones to silty sandstones, massive or with bioturbation, root traces, and clay slickenside. These

facies also contain carbonate nodules (Pc) (Fig. 4D) or hematite nodules (Fh) (Fig. 4E). The contacts
between conglomeratic and silty facies are sharp or in a rapid transition.

In the Valle Encantado section, channel-fill facies exhibit shallow lenticular to lenticular geometries, with common gently dipping strata (FA2). Sediments consist of pinky to whitish, medium- to fine-grained sandstones with trough cross-stratification, climbing ripples, and frequent mud-clasts, arranged in 2-2.5-m-thick fining-up successions (Fig. 4F). The fine-grained floodplain facies are dominated by red, mostly parallel-laminated, sandy to clayey siltstones, with scarce clay slickenside. Contrary to those observed in the western areas, the floodplain facies lack carbonate or hematite nodules.

233 The lateral gradation of the facies associations from western to eastern areas displays channels 234 dominated by bed-load sediments to mixed-load sediments. Furthermore, the rapid transition from 235 coarse-grained channels and bioturbated tops observed in Luracatao, Tin Tin, and Tonco suggests 236 abandonment of active channels and/or avulsion processes common in braided fluvial systems (Miall, 237 1996). Toward the Valle Encantado site, lenticular channels with common inclined strata geometries 238 are dominated by mixed-load sediments, displaying a grading transition to overbank deposits. The 239 facies association is consistent with channel belts of moderate to high-sinuosity rivers associated 240 with cohesive floodplains, as previously interpreted (del Papa, 2006).

Recently, Andrews et al. (2017) studied the paleosols of this unit at the Tin Tin and Obelisco sections, indicating a predominance of red claystones with abundant calcic nodules and vertic structures and occasional gray mottles. They interpreted these paleosols as having formed in a relatively dry paleoenvironment and classified them as calcic vertisols in accordance with Mack et al. (1993).

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#### 250 **5. MINERALOGY: X-ray diffraction (XRD)**

Eighty-seven whole rocks and clay sub-samples (32 from Valle Encantado, 18 from Tin Tin, 18 from Tonco, 12 from Luracatao and seven from Obelisco) from the Lumbrera Fm were analyzed by XRD. The stratigraphic position of the samples corresponding to the Valle Encantado, Tonco, Tin Tin, and Luracatao sections are shown in Figures 2 and 3. The results are summarized in Tables 2 and 3.

# 255 5.1 Whole rock mineralogy

256 XRD analysis of bulk samples of the claystones, siltstones, and fine-grained sandstones shows 257 quartz and phyllosilicates to be the major components in association with minor plagioclase, K 258 feldspar, and carbonates. Plagioclase contents are low in all the samples, usually < 10%; likewise K-259 feldspar contents range from 2 to 7%. Calcite is absent in most of the samples or represents less than 260 5%, and rarely reaches 10%. The exceptions are several levels from the Tin Tin section with calcite 261 contents from 25 to  $\sim$  40 %, one bed from the Tonco section displaying over 50% calcite, and one 262 level from Luracatao section showing  $\sim 20\%$  calcite. Hematite is frequently present in subordinate 263 amounts in beds of the Lumbrera Fm from all the studied sections.

## 264 5.2 Clay Mineralogy

265 The clay mineralogy of the samples from the Valle Encantado, Tonco, Obelisco, Tin Tin, and 266 Luracatao sections is summarized in Tables 2 and 3, and representative XRD patterns are shown in 267 Figure 5. The clay mineral assemblages identified in the Lumbrera Fm are commonly dominated by 268 illite-mica or kaolinite (Figs. 5A, B), and less frequently by smectite (Fig. 5C) or illite/smectite mixed 269 layers (I/Sm). The siltstones and fine-grained sandstones from the Luracatao section represent the 270 latter case because they contain abundant smectite and an expandable phase that according to XRD 271 corresponds to R0-type I/Sm with 10 to 30% smectite layers, but whose composition, as will be 272 discussed in the next section, covers a range from R1 and R0-type I/Sm to smectite. For example, the 273 XRD trace of the air-dried oriented mount of LUR-5 displays the characteristic reflections of illite-

muscovite and kaolinite, but also a peak at 14.98 Å, whereas in the ethylene-glycol-(EG)-treated pattern this peak shifts to 16.74 Å, and reflections at 8.54 and 5.57 Å also appear, indicating RO-I/Sm (Fig. 5D). Chlorite seldom occurs as a subordinate phase in any of the sections. Substantial changes in clay mineral assemblages are observed from bottom to top in each section but also between the different sections (Figs. 2, 3).

279 In the Valle Encantado site, large variations in the relative abundances of kaolinite (6-79%) and 280 illite-mica (19-91%) are observed, as indicated by the Kln/Ms ratio varying from 0.07 to 4.15. 281 Moreover, I/Sm or rarely smectite occurs in subordinate amounts in the clay fraction (0-15%), except 282 in the upper  $\sim$  50 m of the section, where both phases are absent (Table 2, Fig. 2A). The fluctuations 283 in the relative abundances of kaolinite and illite-mica along the stratigraphic section are in some 284 cases gradual and in others abrupt. Indeed, the basal level of the Lumbrera Fm has a KIn/Ms ratio of 285 0.68, which is followed in a bed 7 m up-section by a sharp increase in kaolinite content, coupled with 286 an abrupt decrease in the illite-mica percentage, giving rise to a KIn/Ms ratio of 4.15 (Fig. 2A, Table 287 2). Subsequently, the illite-mica relative abundances increase sharply at the expense of kaolinite in a 288 siltstone level 3 m up-section, displaying a Kln/Ms ratio of 0.79; after this decrease, a shale bed 4 m 289 upwards shows a KIn/Ms ratio of 1.49. The KIn/Ms ratio remains low over the next  $\sim$  50 m and 290 thereupon increases sharply from 0.19 to 1.03 in 17 m. Towards the top of the unit, three more 291 segments of low KIn/Ms followed by sharp increases in the relative abundances of kaolinite are 292 observed, displaying maximum Kln/Ms ratios of 1.54, 0.94 and 2.12, respectively (Table 2, Fig. 2A).

At the Tonco site, where illite is the most abundant clay mineral, the kaolinite content also shows ups and downs, but it remains below 40% throughout the section (Table 2, Fig. 2B). Consequently, changes in the relative abundance of kaolinite and the Kln/Ms ratio (0.16 to 0.84) are less marked than in the Valle Encantado section. Moreover, in this section I/Sm, or less frequently smectite, represents a subordinate component of the clay fraction, attaining maximum contents of 35 and 18%, respectively.

The clay assemblages identified along the Obelisco section show a similar pattern to those at the Tonco site, located northward at a similar longitude. In this case, the Kln/Ms ratio varies from 0.14 to 1.27, whereas smectite and less frequently I/Sm occur as subordinate phases (Table 3).

The clay assemblages identified at the Tin Tin section are quite different (Table 3); illite-mica is the dominant phase in most of the beds whereas kaolinite is absent in several levels or occurs in low amounts (5-19%), thus yielding Kln/Ms ratios from 0 to 0.2, except for one level with a higher value (0.64) (Fig. 3A). Furthermore, smectite is present in all the levels analyzed, representing the main phase in two cases (LUT 29a and LUT 37). Noteworthy is that the claystones and siltstones of the upper ~ 49 m of the section contain I/Sm (14-22%) in addition to smectite.

In the Luracatao site, the clay assemblages resemble those of the Tin Tin site in that kaolinite represents a subordinate phase all along the section (6-17%) (Table 3). In this site, however, the clay assemblage is dominated either by illite-mica or by I/Sm (R0)-Sm, as these vary from 31 to 76% and from 10 to 62 %, respectively (Fig. 3B).

In summary, Luracatao and Tin Tin have higher contents of Sm (+ I/S) than of kaolinite, whereas
Valle Encantado, Tonco, and Obelisco have higher abundances of kaolinite than of Sm (+ I/S).

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# 315 6. TEXTURAL STUDY: Scanning electron microscopy

The five samples from the Valle Encantado section (acronym LUVE) chosen for the SEM study show heterometric and heterogeneous textures with abundant detrital fragments mainly composed of quartz, K feldspar, plagioclase, and micas (muscovite and phengite) (Fig. 6A). The composition of plagioclase is Ab<sub>0.7-0.8</sub>, whereas K feldspar displays low Na contents, indicating a probable volcanic origin. The degree of alteration of the feldspars and micas is variable, from slight to moderate in some samples (Fig. 6 A, B) to strong in others (Fig. 6D, E), with the kaolinite contents increasing accordingly. The BSE images show that the kaolinite has grown between the mica sheets (Fig. 6B, D),

sometimes showing displacive precipitation along the cleavage planes, with the feldspars slightly altered, sometimes displaying corroded edges. By contrast, in the samples showing higher kaolinite contents, feldspar (both K feldspar and albite) displays corroded edges and also replacement by kaolinite (Fig. 6D, E), whereas booklets of kaolinite are common in the matrix (Fig. 6E, F). At the same time, the replacement of mica sheets by kaolinite may be partial to almost total, since in some cases only relicts of mica remain within kaolinite booklets. These rocks also show a fine-grained matrix, which according to the EDS analysis consists of K-Al-Si-rich clays, probably of illitic phase (Fig. 6A, C).

330 The six samples from the Tonco section (acronym LULE) and the Obelisco section (acronym LU) 331 studied by SEM show a similar mineralogy and texture to those from Valle Encantado (Fig. 7A). 332 Smectite is more abundant in samples from the Obelisco section than in those from the Tonco 333 section. In this case the alteration of feldspars and micas varies from slight to moderate. In samples 334 with slight alteration, the BSE images show that kaolinite partly replaces large mica laths, with 335 kaolinite growth originating at grain edges leading to the characteristic fanned-out textures (Fig. 7B, 336 C) (De Ros, 1998; Arostegui et al., 2001). Conversely, in the samples showing stronger alteration, 337 kaolinite not only replaces mica but also partially replaces albite and occurs as booklets in the matrix (Fig. 7D). Fig. 7E shows the growth of Sm from altered K feldspar. 338

339 The seven samples from the Tin Tin section (acronym LUT) studied by SEM also show 340 heterometric and heterogeneous textures with abundant detrital fragments mainly composed of 341 quartz, K feldspar, albite, and micas (muscovite and phengite). In general, feldspars and quartz are 342 scarcely altered, displaying net edges, although feldspars showing incipient alteration to smectite 343 occasionally occur. The K feldspars usually have some Na contents (Si<sub>3.0</sub>Al<sub>0.9</sub>K<sub>1.1</sub>Na<sub>0.1</sub>), as is typical of 344 high-temperature feldspars, thus indicating a probable volcanic origin. The occurrence of micron-345 sized idiomorphic feldspars in several samples also suggests some volcaniclastic input. In agreement 346 with the XRD data, SEM observations show that illite is the main clay mineral in most of the samples, 347 followed by smectite in varying proportions, whereas kaolinite is scarce or absent. Textural features

348 indicate that illite-mica is detrital in origin, whereas smectite displays a rose-like texture in fresh cut 349 (Fig. 7F), thus indicating an authigenic origin. Furthermore, BSE images show that kaolinite has grown 350 between the mica sheets and also partially replaced the mica, giving rise to muscovite-kaolinite 351 intergrowths. By contrast, in the samples where XRD shows clay assemblages dominated by smectite 352 (LUT29A and LUT37), BSE images indicate that this smectite occurs in the matrix, or filling cracks in 353 sub-idiomorphic quartz (Fig 7G), with feldspar crystals without evidence of alteration. If smectite 354 does not derive substantially from the feldspars, which are scarcely altered, it likely formed through 355 alteration of a more reactive material contained in the matrix as well as in these cracks, probably 356 volcanic glass.

357 The sample from Luracatao (acronym LUR) studied at the SEM scale displays a heterogeneous 358 texture with abundant detrital fragments, mainly quartz, K feldspar, plagioclase, and micas 359 (muscovite and phengite), whereas the matrix is composed of clay minerals and calcite. The feldspars 360 and quartz display net edges, with no evidence of alteration, whereas mica laths are seldom replaced 361 by kaolinite or smectite along the edges. Morphologies typical of authigenic smectite (or I/Sm) are 362 observed in the matrix and fully replace irregular fragments 20-30 µm long. The coexistence of fresh 363 K feldspar fragments with abundant smectite in the matrix, along with irregular fragments totally 364 replaced by smectite-I/Sm, suggests (Fig. 7H) that this would have formed from a highly reactive 365 material, probably volcanic glass, all the more so as these fragments resemble glass shards. 366 According to EDS analysis, the composition of these dioctahedral clays covers the range from R1-type 367 I/Sm to dioctahedral smectite.

In general, the SEM study shows that the matrix of the analyzed rocks is composed of mixtures of dioctahedral clays such as illites, I/S, and smectite. Overall, the detrital silicates show a variable degree of alteration. The formation of smectite and kaolinite has been described in the SEM study from Al-K-silicates, whereas kaolinite also grows between the mica sheets. Therefore, smectite and kaolinite are authigenic and were formed by weathering processes. The presence in Luracatao of

irregular fragments formed by smectite-type clays, along with the smectite-rich matrix in samples with slightly altered silicates (e.g., feldspars), suggests the crystallization of smectite from vitreous volcanic material. Illitic phases such as illite and mica would be detrital phases that resisted the weathering. The origin of the I/S phases might be a consequence of the alteration of illitic phases, or alternatively they may have formed from transformations of the volcanic smectite.

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# 379 7. CHEMICAL COMPOSITION OF SMECTITE-TYPE CLAYS: X-ray dispersive analysis (EDS)

380 Smectites can have variable compositions depending on their genesis, unlike kaolinite, which 381 usually has a fixed composition. For this reason, the compositional study of clays focused on 382 smectite-type clays.

383 All the smectite-type clays analyzed (smectite to R1-type I/Sm) correspond to the Al-rich-384 dioctahedral subgroup, with K>Ca>Na in interlayer sites (Table 4). As can be seen in Figure 8A, the 385 smectites from the Luracatao section show Fe/Fe+Mg << 0.50 and are more siliceous than those from 386 Obelisco, Tin Tin, Tonco, and Valle Encantado, which display Fe/Fe+Mg > 0.50. On the other hand, 387 the smectite-type clays from the Tin Tin section display a wider compositional range. The Si vs. Mg 388 plot was employed to differentiate beidellite from montmorillonite. As can be seen in Figure 8B, the 389 smectite-type clays from Luracatao mainly correspond to beidellite, whereas in siltstones and 390 claystones from Tonco, Valle Encantado, and Obelisco montmorillonite prevails. The smectite-type 391 clays from the Tin Tin section show a wide variation covering both fields. It is worth noting that 392 smectites with lower Si contents and higher Fe/Fe+Mg ratios prevail in the sites in which kaolinite is 393 more abundant than smectite. By contrast, smectites that show higher Si contents and lower 394 Fe/Fe+Mg ratios correspond to sites in which smectite is more abundant than kaolinite. These 395 compositional differences suggest that smectite-type clays from Luracatao and some of those from 396 the Tin Tin section were formed from a different parent material than those from Valle Encantado, 397 Tonco, and Obelisco.

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# 399 8. DISCUSSION

400 Burial diagenesis can obliterate or even destroy the paleoclimatic signal recorded in clay mineral 401 assemblages, so this is the first variable to be evaluated. In our case the occurrence of smectite in the 402 basal levels of the Lumbrera Fm in the Tin Tin, Tonco, and Obelisco sections implies that these 403 sediments were only affected by early diagenesis during their burial history. This is in agreement with 404 previous studies indicating that the underlying Maíz Gordo Fm only underwent eogenetic diagenesis 405 (Do Campo et al., 2007, 2018). However, I/Sm mixed layers occur in most of the levels of Valle 406 Encantado, Tonco, and Luracatao, in the uppermost levels of the Tin Tin section, and in some levels 407 of the Obelisco section. The vertical distribution of I/Sm along each stratigraphic column is 408 inconsistent with an origin by burial diagenesis (Figs. 2 and 3). Moreover, most of the beds that 409 contain I/Sm display a KIn/Ms ratio less than 1, implying low to moderate weathering and suggesting 410 that I/Sm could have been formed by moderate chemical weathering of illitic phases, or by early 411 alteration of highly reactive volcanic glass. The textures observed in the BSE images point to a 412 volcaniclastic origin, as was noted in the previous section. Kaolinite and smectite present in these 413 sediments are primary in origin, and thus the lateral and vertical changes in the clay mineral 414 assemblages may result from changes in paleoclimate and weathering regimes, which are the first-415 order factors controlling sediment accumulation. Although syn-sedimentary tectonic and volcanic 416 events can also leave a recognizable fingerprint on clay mineral assemblages, the deposition of the 417 Lumbrera Fm corresponds with a period of tectonic calm, so only volcanic input remains as an 418 additional variable to be considered.

Based on varying paleosol types, organic carbon isotope signatures, and chemical weathering indexes, Andrews et al. (2017) inferred a marked change in annual mean temperatures and precipitation (MAP) between the top of the underlying Maíz Gordo Fm and the basal levels of the Lumbrera Fm. Moreover, they correlated the CIEs registered in the top 30 m of the Maíz Gordo Fm

423 with the PETM. As regards clay mineral assemblages, the basal levels of the Lumbrera Fm display a 424 sharp decrease in kaolinite contents and in the KIn/Ms ratio in comparison with the upper levels of 425 the Maíz Gordo Fm (Do Campo et al., 2018). Indeed, in the Tin Tin section the upper  $\sim 40$  m of the 426 Maíz Gordo Fm showed an average kaolinite content of 63%, with peak values of 82%, contrasting sharply with 7% in the basal level of the Lumbrera Fm in the same section, but also with the 427 428 percentage observed in other sections (8 to 39%). This change in the clay mineral assemblages, 429 already observed by Do Campo et al. (2018), was interpreted as resulting from a strong paleoclimatic 430 change, in agreement with Andrews et al. (2017).

431 However, the lateral changes in clay mineral assemblages observed in the Lumbrera Fm could 432 not have arisen from changes in climatic conditions, as at present the Valle Encantado and Tin Tin 433 sites are located at a distance of only ~24 km, whereas the Luracatao and Tin Tin sites are ~36 km 434 apart (Fig. 1). What is more, the fluvial styles are consistent with downslope variation from west to 435 east, and there are no significant changes in sedimentary facies that might suggest contrasting 436 precipitation regimes. According to their clay mineralogy, a first distinction can be made between 437 eastern (Valle Encantado, Tonco, and Obelisco) and western sections (Tin Tin and Luracatao). In the 438 first group one third to half of the beds, depending on the section, display kaolinite contents over 439 30%, in some cases representing the dominant phase in the clay assemblages.

440 In the western sections, by contrast, kaolinite is absent or represents less than 10% in most of 441 the beds, or occasionally as much as 19%. What is more, in the Tin Tin section two levels display 442 smectite contents higher than 50% (LUT-29a, LUT-37), in contrast with the clay mineralogy of the 443 adjacent beds. At the SEM scale, evidence of smectite (or I/Sm) that originated from a parent 444 material more reactive than crystalline feldspars or quartz was observed in samples from the Tin Tin 445 and Luracatao sites. Namely, the occurrence of smectite filling cracks in quartz fragments showing no 446 evidence of alteration, irregular fragments resembling glass shards totally replaced by smectite-I/Sm (Fig. 7G), as well as abundant smectite in siltstone levels containing fresh feldspar fragments, 447

448 suggests that smectite-type clays must have formed from a highly reactive material, probably 449 volcanic glass. K feldspar with minor Na contents, as well as the occurrence of euhedral K feldspar 450 grains, indeed point to volcaniclastic input in the sediments in the western sections. Consequently, 451 the smectite-type clays forming these sediments probably derived not only from weathering of 452 silicates during edaphization processes, but also from early alteration of glassy volcaniclastic 453 material. Therefore, the clay mineral assemblages of the western sections of the Lumbrera Fm are 454 not suitable for drawing paleoclimatic inferences as the climatic signal has been obliterated by the 455 volcaniclastic input. The distribution of smectite and kaolinite in the area may be a consequence of 456 the heterogeneous scattering of the volcanic material both in vitreous and crystalline phases.

The fact that the smectites originated from different material precursors is reflected in their compositions. Smectite-type clays from Luracatao mainly correspond to beidellite, whereas in the siltstones and claystones from Valle Encantado and Obelisco montmorillonite prevails, and smectitetype clays from the Tin Tin section show a wide variation covering both fields. Smectites displaying lower Si contents and higher Fe/Fe+Mg ratios prevail in the sections in which kaolinite is more abundant than smectite. By contrast, smectites that show higher Si contents and lower Fe /Fe+Mg ratios correspond to sections in which smectite is more abundant than kaolinite.

SEM images revealed evidence of volcaniclastic material in samples from Luracatao, so this could be the main precursor of the smectite-type clays in this section. On the other hand, smectitetype clays from Valle Encantado, Tonco, and Obelisco must have formed from the weathering of the detrital aluminum silicates. In both cases, the smectites would be authigenic and would be the consequence of *in situ* weathering.

The wide compositional range of smectite-type clays from the Tin Tin section suggests that they correspond to a blend of smectites of different origins both through the alteration of volcanic constituents and weathering of silicates.

472 On the other hand, the siltstones and claystones from the eastern sections exhibit varying 473 degrees of alteration in feldspars and micas, ranging from slight to moderate in Tonco and Obelisco 474 to intense in the case of the Valle Encantado section. A higher degree of weathering in detrital 475 minerals is consistent with higher kaolinite abundance. Furthermore, textural evidence indicates that 476 the kaolinite is authigenic in origin, as BSE images of the rocks with moderate to strong weathering 477 show feldspar fragments with corroded edges and also replacement by kaolinite, whereas booklets 478 of kaolinite are common in the matrix. Indeed, BSE images reveal partial to almost total replacement 479 of mica sheets by kaolinite, which sometimes shows displacive precipitation along the cleavage 480 planes, whereas in some cases only relicts of mica remain within kaolinite booklets. Therefore, the 481 vertical changes in the relative abundance of clay minerals observed in the eastern sections, essentially the repeated peaks in the kaolinite percentage and Kln/Ms, indicate stages of enhanced 482 483 chemical weathering probably driven by paleoclimatic changes. This is most marked throughout the 484 Valle Encantado section, which represents the most complete profile of the Lumbrera Fm, as six 485 levels display a KIn/Ms ratio  $\geq$  1, in sharp contrast with the clay mineralogy of the rest of the beds 486 (Fig. 2A). These levels probably correspond to periods of humid-subtropical to tropical climates 487 because such are the conditions that favor kaolinite formation under near-surface/meteoric 488 environments at a regional scale (Chamley, 1989; Hallam et al., 1991; Righi and Meunier, 1995; 489 Ruffell et al., 2002). By contrast, the levels in between displaying abundant illite-mica (51-93%) probably formed under a relatively dry paleoclimate. Andrews et al. (2017) carried out carbon 490 491 isotope analysis of organic matter from paleosol levels of the Lumbrera Fm and identified four CIEs in 492 this unit, which they correlated with hyperthermal events occurring during the deposition of the 493 Lumbrera Fm. The CIEs identified by Andrews et al. (2017) correlate quite well with the four beds 494 from the base to the middle of the Valle Encantado section displaying a Kln/Ms ratio  $\geq$  1. Conversely, 495 the intermediate bed showing a KIn/Ms ratio of 0.94 cannot be correlated with the CIEs recognized 496 by these authors. Moreover, the bed from the top of the Lumbrera Fm showing a KIn/Ms of 2.12

497 cannot be correlated with the data from Andrews et al. (2017), as the section considered by these
498 authors has a thickness of ~ 144 m, whereas in the Valle Encantado section it is 231 m thick.

Given the lack of absolute ages for the Lumbrera Fm, the correlation of the six hyperthermal events recorded in this succession with those identified in deep-sea cores can only be tentative. However, there are some age constraints that may be valuable. In the first place, the PETM was identified in the upper part of the underlying Maíz Gordo Fm (Andrews et al., 2017). Second, a middle-late Eocene age has been proposed for the overlying Quebrada de Los Colorados Fm based on mammal fossils and the U-Pb ages of detrital zircons (Payrola et al., 2009; Carrapa et al., 2012). In this scenario a Ypresian to probably early Lutetian age could be postulated for the Lumbrera Fm.

506 The PETM, recognized at the top of the Maíz Gordo Fm, is the most drastic hyperthermal 507 episode of the Paleogene, with an increase in sea temperature of possibly up to 6 °C (Zachos et al., 508 2008). Subsequently, temperatures went back to their pre-PETM levels, before starting a gradual 509 long-term warming trend that reached its maximum during the Early Eocene Climatic Optimum 510 (EECO; ~53-49 Ma; Zachos et al., 2001, 2008; Kirtland Turner and Ridgwell, 2013; Laurentano et al., 511 2015, 2018; Westerhold et al., 2018; Crouch et al., 2020). This trend was punctuated by cyclic 512 hyperthermals throughout the Ypresian stage, ETM2 and 3 being the most severe of these events, 513 initially identified in early Eocene marine sediments (Röhl et al., 2005; Zachos et al., 2005, 2010). 514 Afterwards, high-resolution studies of deep-sea cores identified several prominent CIEs that took 515 place between 56 and 52 Ma and were labeled by Cramer et al. (2003) from E to L (Cramer et al., 516 2009; Zachos et al., 2010; Laurentano et al., 2015, 2018; Thomas et al., 2018; Westerhold, et al., 517 2018; Crouch et al., 2020). Subsequently, 20 additional smaller CIEs have been documented for the 518 Ypresian to early Lutetian (56–45 Ma) based on stable isotope records from sea cores from the 519 Demerara Rise in the equatorial Atlantic (Kirtland Turner and Ridgwell, 2013; Sexton et al., 2011). 520 According to Crouch et al. (2020), the base of the EECO is event J (52.8 Ma, Littler et al., 2014), 521 whereas its top has been linked with a CIE in the uppermost chron C22n dated at 49.1 Ma. This is in

522 agreement with Westerhold et al. (2018), who consider that the EECO lasted 4.12 Ma. The 523 hyperthermal ETM3 took place around 52.5 Ma, thus during the EECO, and temperatures were probably higher than during ETM2 (Thomas et al., 2018). The end of the EECO coincides with the 524 onset of a long-term cooling trend evident in the benthic  $\delta^{18}$ O record (Westerhold et al., 2018). 525 Furthermore, early Eocene hyperthermal events probably triggered increases in precipitation and 526 527 thus intensification of the hydrological cycle, as the amount and intensity of rainfall are controlled by 528 global temperatures (Held and Soden, 2006; Bataille et al., 2016). Accordingly, hyperthermal events 529 should be associated with enhanced chemical weathering due to the combined effects of higher CO<sub>2</sub> levels, which favor silicate alteration, and higher humidity, leading to stronger lixiviation in 530 531 weathering profiles.

532 In light of the previous data for the Maíz Gordo Fm (Andrews et al., 2017; Do Campo et al., 533 2018) and the large body of evidence from the study of deep-sea cores, we tentatively correlate the 534 basal bed of the Lumbrera Fm at the Valle Encantado section, which displays a Kln/Ms ratio of 4.15, 535 with ETM2, and the middle level, which has a KIn/Ms ratio of 1.54, with ETM3. According to the 536 relative abundance of kaolinite and the KIn/Ms ratio, chemical weathering was markedly stronger 537 during ETM2 than during ETM3, implying that in this inner continental basin the paleoclimate was more humid and warmer during ETM2, in contrast with the evidence from marine sediments 538 (Thomas et al., 2018). On the other hand, the clay mineral assemblage from the top bed of this 539 540 section, which shows a KIn/Ms ratio of 2.12, is in marked contrast with the mineralogy of the basal 541 levels of the overlying Quebrada de Los Colorados Fm, which are dominated by illite-mica with 542 subordinate smectite, with kaolinite representing less than 10%. Based on the clay mineral 543 assemblages, temperate semi-arid to arid paleoclimatic conditions have been inferred for this unit 544 (Do Campo et al., 2010). This switch in paleoclimate from subtropical humid to warm and arid would 545 thus correspond with the global changes registered in the transition from the Ypresian to Lutetian stages. Accordingly, the upper level of the Lumbrera Fm corresponds with the peak of the EECO, 546

which was followed by a long period of deep-sea cooling (Chekar et al., 2018). The other two levels displaying a Kln/Ms  $\geq$  1 correspond to some of the minor hyperthermal events identified in marine records. Nonetheless, absolute ages would be required for more precise correlations.

In the Tonco section the Kln/Ms ratio also undergoes ups and downs, but the changes are less pronounced. Several beds from the base, the middle, and the top of the section display a Kln/Ms > 0.6 in contrast with the surrounding levels. The two beds from the base with Kln/Ms ~ 0.7 could be correlated with ETM2, and the bed from the top showing a Kln/Ms of 0.84 could be associated with ETM3, whereas the intermediate bed with a Kln/Ms of ~ 0.7 could correspond to one of the intermediate hyperthermal events recorded in the Valle Encantado section.

At the Obelisco site the sampling was less detailed than in the other sections, so only a few fluctuations in kaolinite percentages can be observed. Indeed, there is no evidence of ETM2, and only one level from the top of the section with a Kln/Ms ratio of 1.27 could be tentatively correlated with ETM3.

560 To shed light on the impact of past global greenhouse conditions on continental subtropical 561 areas, Kelson et al. (2018) performed a study of the Tornillo Basin in Texas (USA), covering terrestrial 562 sediments from the Paleocene to early Eocene (67 - 52 Ma, chron C30n to chron C23r). The authors 563 performed clumped isotope analysis on pedogenic calcite nodules and determined that the average 564 temperatures from the early Eocene were higher than those from the Paleocene  $(32 \pm 2 \text{ and } 25 \pm 3)$ 565 °C, respectively; Kelson et al., 2018), with the caveat that they excluded the PETM from their 566 statistical analysis. The authors consider that this shift in temperatures at the start of the Eocene 567 most likely relates to a contemporaneous gradual rise in the concentration of atmospheric pCO<sub>2</sub> 568 (Kelson et al., 2018). According to the same authors, the Paleocene environment in the Tornillo Basin 569 was probably subtropical and humid with year-round precipitation, whereas for the Eocene the 570 climate models as well as sandstone sedimentology suggest more seasonal precipitation with stronger summer monsoon rainfall than during the Paleocene (Bataille et al., 2018). In the case of the 571

Salta Basin, several proxies including climatic transfer functions applied to paleosol levels of the Maíz Gordo Fm, along with  $\delta^{13}C_{org}$  (Andrews et al., 2017) as well as clay mineral assemblages (Do Campo et al., 2018), indicate subtropical and humid paleoclimatic conditions during the upper Paleocene, in agreement with the studies of the Tornillo Basin. Clay mineral assemblages also suggest contrasting climatic condition between the Maíz Gordo and Lumbrera formations, as these exhibit average kaolinite contents of 52% and 31%, respectively. However, our data do not indicate warmer conditions for the early Eocene, as suggested for the Tornillo Basin.

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## 580 9. CONCLUSIONS

The study of the clay mineralogy of the fluvial detrital rocks of the Lumbrera Fm in the western part of the Salta Basin (NW Argentina) shows the coexistence of detrital clays (illitic phases) and authigenic clays (smectite to R1-type I/Sm and kaolinite).

The authigenic clays were formed from the alteration of detrital crystalline fragments (feldspars and micas) and from the alteration of volcanic glass, which is much more reactive than the crystalline phases. The presence of different precursors for the smectite-type clays favors variable compositions, from beidellite to montmorillonite. The variable distribution of smectite and kaolinite between the different sites and the differences in smectite composition suggest a heterogeneous scattering of the volcanic material both in vitreous and crystalline phases over the study area.

This study thus shows that the relative proportions of smectite and kaolinite observed in the Tin Tin and Luracatao sites are not indicative of variable climatic conditions, since the crystallization of smectite-type clays from volcanic glass obliterates the climatic signal. On the other hand, in the case of the Valle Encantado site, where smectite-type clays are almost absent, the input of volcanic material can be inferred to have been very low (or absent). Thus, the kaolinite crystallization in this site suggests the presence of periods of humid-subtropical to tropical climate, because such conditions favor kaolinite formation under near-surface/meteoric environments at a regional scale.

The increase in the Kln/Ms ratio (4.15) in the basal bed of the Lumbrera Fm at Valle Encantado suggests the presence of cyclic hyperthermals through the Ypresian stage and can be correlated with ETM2; the Kln/Ms ratio (1.54) in the middle level can be correlated with ETM3. These are the major hyperthermal events identified in early Eocene marine sediments.

This research shows that clay mineralogy is a powerful paleoclimatic proxy in terrestrial sedimentsnot affected by diagenetic changes.

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# 831 Figure and Table captions

Figure 1. Satellite image of northwestern Argentina with rectangles outlining the location of geological maps and black stars showing the location of stratigraphic sections. A) Geological map of the Valle Encantado, Tin Tin and Tonco areas. B) Geological map of the Luracatao area. C) Geological map of the Obelisco area. Geological maps A and C were modified from Vergani and Starck (1989), and B from Payrola Bosio et al. (2009).

Figure 2. Measured stratigraphic sections and clay mineralogy of the fine-grained fraction based on XRD results. A) Valle Encantado site. B) Tonco-La Escalera site. Levels displaying peak Kln/Ms ratios mentioned in the text are indicated with arrows.

Figure 3. Measured stratigraphic sections and clay mineralogy of the fine-grained fraction based on XRD results. A) Tin Tin site. The arrow indicates a bed displaying the highest Kln/Ms ratio, further details in the text. B) La Represa-Luracatao site. Figure 4. Field photographs of the Lumbrera Fm showing characteristic aspects of channel-fill facies and fine-grained floodplain facies at (A) Luracatao; solid square detail shown in (D), and (B) Tonco, note the person for scale. C) Conglomerate grading to sandstones with trough cross-stratification (Gt-St facies). D) Close-up view of a calcic-rich paleosol level (Pc facies) interbedded in floodplain fines at Luracatao. E) Grey to yellowish mudstone level denoting hematite formation (Ph facies). F) Sandy channel-fill facies displaying normal gradation to overbank sedimentation. Arrow indicates mudstone level sampled, Valle Encantado.

Figure 5. XRD patterns of air-dried and ethylene-glycol-solvated clay fractions of representative
samples. A) LUVE27. B) LUVE32. C) LUT37. D) LUR5.

Figure 6. SEM/BSE images of samples from Valle Encantado. A) Textural image (LUVE-13) showing heterometric texture with detrital Qtz, KF, Ms, and fine-grained clay matrix. B) KIn-Ms intergrowths and kaolinite (KIn) booklets (LUVE-2). C) Clay-rich matrix and detrital KF, Qtz and Ms. D) KIn/Ms intergrowths, altered KFand Qtz fragments. E) and F). Authigenic KIn booklets and altered Ab and KF.

Figure 7. SEM/BSE images. A) Textural image (LULE-16) showing heterometric texture with detrital fragments of Qtz, KF, Ab, Ms and fine-grained clay matrix. B) Sample LULE-25 with KIn-Ms intergrowths and sparitic calcite cementing the rock. C-D) Abundant detrital fragments of silicates and KIn-Ms intergrowths (sample Lu9a). E) Growth of smectite from altered KF in sample LU29a. F) Smectite flakes forming the clay matrix of LUT39. G) Smectite filling cracks in quartz in sample LUT29a. H). Anhedral fragments formed by smectite flakes.

Fig. 8. SEM/EDS analyses of smectites. A) Fe/Fe+Mg vs Si (apfu). B) Mg (apfu) vs. Si (apfu).

Table 1. Facies association identified in the Lumbrera Formation. Codes from Miall (1996).

Table 2. Clay mineralogy of samples from the Valle Encantado and Tonco sites based on XRD analyses. Kln/Ms values highlighted in grey correspond to peak values mentioned in the text and marked with arrows in Figure 2.

- Table 3. Clay mineralogy of samples from the Obelisco, Tin Tin, and Luracatao sites based on XRD analyses. Kln/Ms values highlighted in grey correspond to peak values mentioned in the text and marked with an arrow in Figure 3.
- Table 4. Average composition of smectite-type clays (smectite to R1-type I/Sm) according to EDS
- 871 microanalyses. Σoct: Sum of octahedral cations, F/FM= Fe/(Fe+Mg).

















|     | ARCHITECTURAL<br>ELEMENTS         | FACIES                    | FEATURES   | INTERPRETATION   |  |  |  |
|-----|-----------------------------------|---------------------------|--|--|--|--|--|
| FA1 | CH, GB-SB                         | Gt, Gp, St,<br>Sp, Sm, Fm | White to pinkish, medium to fine-granied<br>conglomerates to gravelly sandstones.<br>Shalow erosive bases, gravel to granule<br>lag, nomal grading, trough cross-<br>stratification, minor planar cross-<br>stratification, imbrication, Bioturbated tops                                    | Dilute currents w ith mainly tractive<br>load. Filling scours, shallow channels,<br>2D-3D bar migration. Overbank<br>sedimentation     |  |  |  |
| FA2 | CH, SB,LA,LV St, Sp, Sr,<br>SI,Sm |                           | Thick to thin lenticular to shallow lenticular<br>strata. Inclined bedding. White to pinky<br>fine to coarse-grained sandstones.<br>Irregular bases, normal grading, trough<br>and planar-cross stratification. Climbing<br>ripples, some contorted beds and load<br>casts, bioturbated tops | Belts of channels filling with sandy<br>material. Mesoforms migration by<br>lateral- and dow nstream- bar acretion.<br>Levee formation |  |  |  |
| FA3 | OF-CR                             | Fm, Fl,Pc,<br>Ph, Sm      | Red, massive, slightly stratified siltstones<br>to sandy siltstones. Carbonate nodules,<br>vertical tubes, mottling. Sheet-like,<br>massive or bioturbated sandstones  | Floodplain sedimentation, calcic-rich paleosols. Crevasse channels and splays  |  |  |  |

Table 1. Facies association identified in the Lumbrera Formation. Codes from Miall (1996).

Table 2

| Location | Location % KIn     |          | llite-Ms | Sm  | l/Sm             | Chl  | Kln/Ms |
|----------|--------------------|----------|----------|-----|------------------|------|--------|
|          | LUVE 32            | 68       | 32       |     |                  |      | 2.12   |
|          | LUVE 31            | 18       | 82       |     |                  |      | 0.22   |
|          | LUVE 30            | 8        | 92       |     |                  |      | 0.09   |
|          | LUVE 29            | 6        | 94       |     |                  |      | 0.07   |
|          | LUVE 28            | 9        | 91       |     |                  |      | 0.10   |
|          | LUVE 27            | 22       | 78       |     |                  |      | 0.29   |
|          | LUVE 26            | 10       | 69       |     | 12               | 9    | 0.14   |
|          | LUVE 25            | 18       | 70       |     | 4                | 8    | 0.26   |
|          | LUVE 23            | 45       | 48       |     | 6                |      | 0.94   |
|          | LUVE 22            | 13       | 72       |     | 15               |      | 0.18   |
|          | LUVE 21            | 28       | 67       |     | 5                |      | 0.42   |
| 0        | LUVE 20            | 59       | 38       | 3   |                  |      | 1.54   |
| ad       | LUVE 18            | 10       | 78       |     | 12               |      | 0.13   |
| ant      | LUVE 17            | 37       | 59       |     | 4                |      | 0.62   |
| 20       | LUVE 16            | 28       | 61       |     | 11               |      | 0.46   |
| ш        | LUVE 15            | 33       | 60       |     | 8                |      | 0.54   |
|          | LUVF 14            | 39       | 57       | 4   |                  |      | 0.70   |
| - Ka     | LUVE 13            | 47       | 46       | •   | 8                |      | 1.03   |
|          | LUVE 12            | 15       | 81       |     | 3                |      | 0.19   |
|          |                    | 16       | 77       |     | 7                |      | 0.10   |
|          |                    | 35       | 58       |     | 7                |      | 0.21   |
|          |                    | 37       | 50       |     | 1                |      | 0.62   |
|          |                    | 18       | 75       |     | - <del>-</del> 7 |      | 0.02   |
|          |                    | 30       | 55       |     | 6                |      | 0.24   |
|          |                    | 34       | 60       |     | 7                |      | 0.72   |
|          |                    | 50       | 20       |     | ו<br>ר           |      | 1.40   |
|          |                    | 12       | 53       |     | 5                |      | 0.70   |
|          |                    | 42<br>70 | 10       |     | 2                |      | 0.79   |
|          |                    | 25       | 13<br>51 |     | 11               |      | 4.15   |
|          | LOVE I             | - 35     | 51       |     | 14               |      | 0.00   |
| Location | 0/                 | Kln      | llito-Me | Sm  | l/Sm             | Chl  | Kln/Me |
| LUCATION | /0<br>        E 30 | 26       | 68       | 6   | <i>V</i> 3III    | CIII | 0.38   |
|          |                    | 20       | 45       | 0   | 19               |      | 0.30   |
|          |                    | 20       | 4J<br>64 |     | 16               |      | 0.04   |
|          |                    | 20       | 70       |     | 10               |      | 0.52   |
|          |                    | 17       | 70<br>59 |     | 4                | 0    | 0.00   |
|          | LULE 27            | 25       | 50       |     | 10               | 9    | 0.29   |
| era      | LULE 20            | 30       | 62       |     | 3<br>6           | 0    | 0.50   |
| ale      | LULE 23            | 32       | 6Z       | 0   | 0                | 0    | 0.52   |
| SC       | LULE 24            | 30       | 55       | 9   | 00               | 40   | 0.00   |
| ы<br>Ш   | LULE 23            | 25       | 30       |     | 29               | 10   | 0.71   |
|          | LULE 22            | 23       | 55       |     | 22               |      | 0.41   |
|          | LULE 21            | 10       | 55       |     | 35               |      | 0.18   |
| ō        | LULE 19            | 21       | 70       |     | 9                |      | 0.30   |
|          | LULE 18            | 13       | /5       | 4.5 | 12               |      | 0.17   |
|          | LULE 16            | 11       | 71       | 18  |                  |      | 0.16   |
|          | LULE-15            | 17       | 72       | 12  |                  |      | 0.24   |
|          | LULE-14            | 13       | 83       | 4   |                  |      | 0.16   |
|          | LULE-13            | 39       | 54       | 6   |                  |      | 0.72   |
|          | LULE-12            | 39       | 57       | 4   |                  |      | 0.69   |

Table 2. Clay mineralogy of samples from the Valle Encantado and Tonco sites based on XRD analyses. Kln/Ms values highlighted in grey correspond to peak values mentioned in the text and marked with arrows in Figure 2.

| Location   | า %     | Kln | llite-Ms | Sm     | I/Sm       | Chl | Kln/Ms |
|------------|---------|-----|----------|--------|------------|-----|--------|
|            | LU 10   | 10  | 70       | 21     |            |     | 0.14   |
|            | LU9c    | 36  | 54       | 10     |            |     | 0.65   |
| SCC        | LU 9b   | 32  | 57       | 12     |            |     | 0.56   |
| eli        | LU 9a   | 46  | 36       | 9      | 8          |     | 1.27   |
| l q        | LU 8b   | 9   | 56       |        | 20         | 15  | 0.16   |
|            | LU 8 a  | 15  | 49       |        | 20         | 16  | 0.31   |
|            | LU7a    | 14  | 80       | 6      |            |     | 0.17   |
|            |         |     |          |        |            |     |        |
| Location   | า %     | Kln | llite-Ms | Sm     | l/Sm       | Chl | Kln/Ms |
|            | LUT43   | 10  | 57       | 14     | 18         |     | 0.18   |
|            | LUT42b  | 6   | 49       | 22     | 14         | 9   | 0.11   |
|            | LUT42   | 8   | 60       | 14     | 18         |     | 0.13   |
|            | LUT 41  | 19  | 30       | 32     | 18         |     | 0.64   |
|            | LUT40   | 8   | 53       | 16     | 22         |     | 0.16   |
|            | LUT 39  | 5   | 88       | 7      |            |     | 0.06   |
|            | LUT 38  | 6   | 69       | 25     |            |     | 0.09   |
|            | LUT 37  | 5   | 36       | 59     |            |     | 0.14   |
| i <b>⊨</b> | LUT 36  | 0   | 87       | 13     |            |     | 0.00   |
| <u>e</u> . | LUT 35  | 0   | 89       | 11     |            |     | 0.00   |
|            | LUT 34  | 0   | 98       | 2      |            |     | 0.00   |
|            | LUT 33  | 0   | 94       | 6      |            |     | 0.00   |
|            | LUT 32  | 13  | 65       | 23     |            |     | 0.20   |
|            | LUT 31  | 16  | 65       | 19     |            |     | 0.25   |
|            | LUT 30  | 9   | 80       | 11     |            |     | 0.11   |
|            | LUT 29c | 11  | 88       | 88 1   |            |     | 0.12   |
|            | LUT 29b | 10  | 81       | 81 9   |            |     | 0.13   |
|            | LUT 29a | 7   | 38       | 55     |            |     | 0.19   |
|            |         |     |          |        |            |     |        |
| Location   | า %     | Kln | llite-Ms | Sm-l/S | Sm-I/Sm R0 |     | Kln/Ms |
|            | LUR12   | 9   | 75       | 10     |            | 6   | 0.12   |
|            | LUR11   | 8   | 64       | 2      | 21         | 7   | 0.13   |
| ess        | LUR10   | 17  | 43       | 4      | 10         |     | 0.39   |
| j d        | LUR9    | 14  | 31       | 5      | 55         |     | 0.46   |
| Re         | LUR8    | 8   | 43       | 4      | 9          |     | 0.18   |
| a          | LUR7    | 10  | 37       | 5      | 53         |     | 0.27   |
|            | LUR6    | 6   | 48       | 46     |            |     | 0.13   |
| ata        | LUR5    | 7   | 31       | 6      | 62         |     | 0.22   |
| ací        | LUR4    | 8   | 69       | 1      | 3          | 10  | 0.11   |
| n.         | LUR3    | 9   | 41       | 5      | 50         |     | 0.22   |
|            | LUR2    | 16  | 58       | 26     |            |     | 0.27   |
|            | LUR1    | 8   | 46       | 4      | 6          |     | 0.17   |

Table 3. Clay mineralogy of samples from the Obelisco, Tin Tin, and Luracatao sites based on XRD analyses. Kln/Ms values highlighted in grey correspond to peak values mentioned in the text and marked with an arrow in Figure 3.

|                    | Si     | AI IV  | AI VI  | Fe     | Mg     | Ti     | К      | Са     | Na     | Int.<br>charge | Soct   | F/FM   |
|--------------------|--------|--------|--------|--------|--------|--------|--------|--------|--------|----------------|--------|--------|
| Obelisco (n=23)    | 3.53   | 0.47   | 1.64   | 0.21   | 0.19   | 0.03   | 0.23   | 0.07   | 0.03   | 0.40           | 2.05   | 0.53   |
| st.dv.             | (0.12) | (0.12) | (0.21) | (0.10) | (0.10) | (0.02) | (0.09) | (0.03) | (0.02) | (0.10)         | (0.05) | (0.05) |
| Tin Tin (n=84)     | 3.73   | 0.27   | 1.35   | 0.32   | 0.31   | 0.03   | 0.27   | 0.11   | 0.03   | 0.52           | 2.01   | 0.49   |
| st.dv.             | (0.15) | (0.15) | (0.16) | (0.14) | (0.07) | (0.02) | (0.09) | (0.03) | (0.03) | (0.07)         | (0.06) | (0.13) |
| Tonco (n=15)       | 3.66   | 0.34   | 1.30   | 0.42   | 0.27   | 0.03   | 0.35   | 0.06   | 0.03   | 0.46           | 2.02   | 0.60   |
| st.dv.             | (0.16) | (0.16) | (0.12) | (0.12) | (0.04) | (0.02) | (0.04) | (0.03) | (0.03) | (0.09)         | (0.05) | (0.06) |
| V.Encantado (n=15) | 3.56   | 0.44   | 1.49   | 0.26   | 0.24   | 0.04   | 0.38   | 0.04   | 0.09   | 0.55           | 2.03   | 0.52   |
| st.dv.             | (0.19) | (0.19) | (0.22) | (0.11) | (0.09) | (0.03) | (0.10) | (0.01) | (0.10) | (0.06)         | (0.07) | (0.09) |
| Luracatao (n=32)   | 3.87   | 0.13   | 1.54   | 0.03   | 0.39   | 0.11   | 0.12   | 0.12   | 0.03   | 0.40           | 2.04   | 0.22   |
| st.dv.             | (0.13) | (0.13) | (0.11) | (0.04) | (0.06) | (0.05) | (0.05) | (0.04) | (0.04) | (0.08)         | (0.04) | (0.08) |

Table 4. Average composition of smectites from the different analysed sites.

# **Declaration of interests**

X The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.
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□The authors declare the following financial interests/personal relationships which may be considered as potential competing interests: