Regular Article - Experimental Physics

Initial performance of the CUORE-0 experiment

D. R. Artusa^{1,2}, F. T. Avignone III¹, O. Azzolini³, M. Balata², T. I. Banks^{2,4,5}, G. Bari⁶, J. Beeman⁷, F. Bellini^{8,9}, A. Bersani¹⁰, M. Biassoni^{11,12}, C. Brofferio^{11,12}, C. Bucci², X. Z. Cai¹³, L. Canonica², X. G. Cao¹³, S. Capelli^{11,12}, L. Carbone¹², L. Cardani^{8,9}, M. Carrettoni^{11,12}, N. Casali², D. Chiesa^{11,12}, N. Chott¹, M. Clemenza^{11,12}, C. Cosmelli^{8,9}, O. Cremonesi^{12,a}, R. J. Creswick¹, I. Dafinei⁹, A. Dally¹⁴, V. Datskov¹², M. M. Deninno⁶, S. Di Domizio^{10,15}, M. L. di Vacri², L. Ejzak¹⁴, D. Q. Fang¹³, H. A. Farach¹, M. Faverzani^{11,12}, G. Fernandes^{10,15}, E. Ferri^{11,12}, F. Ferroni^{8,9}, E. Fiorini^{11,12}, S. J. Freedman^{4,5,b}, B. K. Fujikawa⁵, A. Giachero^{11,12}, L. Gironi^{11,12}, A. Giuliani¹⁶, J. Goett², P. Gorla², C. Gotti^{11,12}, T. D. Gutierrez¹⁷, E. E. Haller^{7,18}, K. Han⁵, K. M. Heeger¹⁹, R. Hennings-Yeomans^{4,5}, H. Z. Huang²⁰, R. Kadel²¹, K. Kazkaz²², G. Keppel³, Yu. G. Kolomensky^{4,21}, Y. L. Li¹³, K. E. Lim¹⁹, X. Liu²⁰, Y. G. Ma¹³, C. Maiano^{11,12}, M. Maino^{11,12}, M. Martinez²³, R. H. Maruyama¹⁹, Y. Mei⁵, N. Moggi⁶, S. Morganti⁹, S. Nisi², C. Nones²⁴, E. B. Norman^{22,25}, A. Nucciotti^{11,12}, T. O'Donnell⁴, F. Orio⁹, D. Orlandi², J. L. Ouellet^{4,5}, M. Pallavicini^{10,15}, V. Palmieri³, L. Pattavina², M. Pavan^{11,12}, M. Pedretti²², G. Pessina¹², V. Pettinacci⁹, G. Piperno^{8,9}, S. Pirro², E. Previtali¹², C. Rosenfeld¹, C. Rusconi¹², E. Sala^{11,12}, S. Sangiorgio²², N. D. Scielzo²², M. Sisti^{11,12}, A. R. Smith²⁶, L. Taffarello²⁷, M. Tenconi¹⁶, F. Terranova^{11,12}, W.D. Tian¹³, C. Tomei⁹, S. Trentalange²⁰, G. Ventura^{28,29}, M. Vignati⁹, B. S. Wang^{22,25}, H. W. Wang¹³, L. Wielgus¹⁴, J. Wilson¹, L. A. Winslow²⁰, T. Wise^{14,19}, L. Zanotti^{11,12}, C. Zarra², B. X. Zhu²⁰, S. Zucchelli^{6,30}

¹ Department of Physics and Astronomy, University of South Carolina, Columbia, SC 29208, USA

² INFN-Laboratori Nazionali del Gran Sasso, Assergi, 67010 L'Aquila, Italy

³ INFN-Laboratori Nazionali di Legnaro, Legnaro, 35020 Padua, Italy

⁴ Department of Physics, University of California, Berkeley, CA 94720, USA

⁵ Nuclear Science Division, Lawrence Berkeley National Laboratory, Berkeley, CA 94720, USA

⁶ INFN-Sezione di Bologna, 40127 Bologna, Italy

⁷ Materials Science Division, Lawrence Berkeley National Laboratory, Berkeley, CA 94720, USA

⁸ Dipartimento di Fisica, Sapienza Università di Roma, 00185 Rome, Italy

⁹ INFN-Sezione di Roma, 00185 Rome, Italy

¹⁰ INFN-Sezione di Genova, 16146 Genoa, Italy

¹¹ Dipartimento di Fisica, Università di Milano-Bicocca, 20126 Milan, Italy

¹² INFN-Sezione di Milano Bicocca, 20126 Milan, Italy

¹³ Shanghai Institute of Applied Physics (Chinese Academy of Sciences), Shanghai 201800, China

¹⁴ Department of Physics, University of Wisconsin, Madison, WI 53706, USA

¹⁵ Dipartimento di Fisica, Università di Genova, 16146 Genoa, Italy

¹⁶ Centre de Spectrométrie Nucléaire et de Spectrométrie de Masse, 91405 Orsay Campus, France

¹⁷ Physics Department, California Polytechnic State University, San Luis Obispo, CA 93407, USA

¹⁸ Department of Materials Science and Engineering, University of California, Berkeley, CA 94720, USA

¹⁹ Department of Physics, Yale University, New Haven, CT 06520, USA

²⁰ Department of Physics and Astronomy, University of California, Los Angeles, CA 90095, USA

²¹ Physics Division, Lawrence Berkeley National Laboratory, Berkeley, CA 94720, USA

²² Lawrence Livermore National Laboratory, Livermore, CA 94550, USA

²³ Laboratorio de Fisica Nuclear y Astroparticulas, Universidad de Zaragoza, 50009 Saragossa, Spain

²⁴ Service de Physique des Particules, CEA/Saclay, 91191 Gif-sur-Yvette, France

²⁵ Department of Nuclear Engineering, University of California, Berkeley, CA 94720, USA

²⁶ EH&S Division, Lawrence Berkeley National Laboratory, Berkeley, CA 94720, USA

²⁷ INFN-Sezione di Padova, 35131 Padua, Italy

²⁸ Dipartimento di Fisica, Università di Firenze, 50125 Florence, Italy

²⁹ INFN-Sezione di Firenze, 50125 Florence, Italy

³⁰ Dipartimento di Fisica, Università di Bologna, 40127 Bologna, Italy

Received: 11 February 2014 / Accepted: 3 July 2014 / Published online: 1 August 2014 © The Author(s) 2014. This article is published with open access at Springerlink.com

Abstract CUORE-0 is a cryogenic detector that uses an array of tellurium dioxide bolometers to search for neutrinoless double-beta decay of ¹³⁰Te. We present the first data analysis with $7.1 \text{ kg} \cdot \text{y}$ of total TeO₂ exposure focusing on background measurements and energy resolution. The background rates in the neutrinoless double-beta decay region of interest (2.47 to 2.57 MeV) and in the α backgrounddominated region (2.70 to 3.90 MeV) have been measured to be 0.071 ± 0.011 and 0.019 ± 0.002 counts/(keV · kg · y), respectively. The latter result represents a factor of 6 improvement from a predecessor experiment, Cuoricino. The results verify our understanding of the background sources in CUORE-0, which is the basis of extrapolations to the full CUORE detector. The obtained energy resolution (full width at half maximum) in the region of interest is 5.7 keV. Based on the measured background rate and energy resolution in the region of interest, CUORE-0 half-life sensitivity is expected to surpass the observed lower bound of Cuoricino with one year of live time.

1 Introduction

Neutrinoless double-beta decay (0vDBD) is a hypothetical lepton number violating process in which two neutrons in an atomic nucleus simultaneously decay to two protons, two electrons, and no electron-antineutrinos: $(A, Z) \rightarrow (A, Z + 2) + 2e^{-}$. Observation of 0ν DBD would establish the Majorana nature of the neutrino, i.e., that the neutrino is its own antiparticle, and may provide insights on the neutrino mass scale and mass hierarchy, depending on 0vDBD rate or rate limit (cf. [1]). The experimental signature for 0vDBD is a peak at the 0vDBD Q-value in the two-electron energy sum spectrum. Several recent experiments have reported new limits on the 0vDBD half-life of ¹³⁶Xe [2,3] and ⁷⁶Ge [4]. For comparison between experiments, half-life limits of different isotopes are usually converted to limits on the effective Majorana mass. This conversion, however, takes into account the phase space factors and nuclear matrix elements, the latter of which introduce large uncertainties from different model calculations [5-11]. The current 0vDBD half-life limit for ¹³⁰Te was set by Cuoricino at 2.8×10^{24} y (90 % C.L.) [12].

CUORE-0 is a cryogenic detector that uses an array of TeO₂ bolometers to search for 0ν DBD in the ¹³⁰Te of the bolometers themselves. ¹³⁰Te is an attractive isotope for a 0ν DBD search because of its relatively high Q-value at 2528 keV [13–15] and its very high natural isotopic abundance at 34.2 % [16]. Cryogenic bolometers measure energy through a rise in the temperature of the detector and have

energy resolutions comparable to high purity Ge detectors: for CUORE-0 style bolometers, the energy resolution (full width at half maximum, FWHM) is typically 0.2% at the 0vDBD Q-value. 0vDBD data taking with CUORE-0 began in March 2013.

CUORE-0 also serves as a technical prototype for CUORE (Cryogenic Underground Observatory of Rare Events) [17], which will consist of 19 towers identical to the single CUORE-0 tower. CUORE-0 is the first tower produced on the CUORE assembly line, and its successful commissioning represents a major milestone towards CUORE. CUORE is in the advanced stages of detector construction at the time of this writing and is scheduled to begin data taking in 2015.

2 CUORE-0 detector

CUORE-0 is a single tower of 52 TeO₂ crystal bolometers operating at a typical base temperature of 13–15 mK in Hall A of the Laboratori Nazionali del Gran Sasso (LNGS) underground facility in Italy. As shown in Fig. 1a, the tower consists of 13 planes of four $5 \times 5 \times 5$ cm³ crystals, held securely inside a copper frame by specially designed polytetrafluoroethylene (PTFE) brackets. The copper frame serves as a thermal bath to cool the crystals through the weak thermal coupling provided by PTFE. Each crystal weighs 750 g, which results in a total detector mass of 39 kg and a total ¹³⁰Te mass of

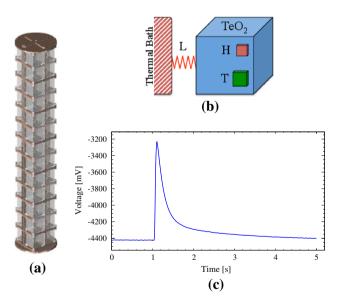


Fig. 1 a CUORE-0 tower array rendering. The tower consists of 13 planes of 4 crystals, mounted in the frame made of copper. **b** Schematic of a single CUORE-0 bolometer showing the thermistor (T), the heater (H), and the weak thermal link (L) between TeO₂ crystal and copper thermal bath (not to scale). **c** An example of a bolometer signal with the energy of approximately 2615 keV. The rise and fall times of this signal are 0.04 s (defined as the time for the pulse amplitude to evolve from 10 to 90% of its maximum) and 0.26 s (90–30% of maximum), respectively

^a e-mail: cuore-spokeperson@lngs.infn.it

^b Deceased

11 kg. Each crystal is instrumented with a single neutron transmutation doped (NTD) germanium thermistor for the signal readout (see Fig. 1b). The typical signal amplitude $\Delta T/\Delta E$ is $10 - 20 \,\mu$ K/MeV. Figure 1c shows an example of a bolometer signal. Additionally, one silicon Joule heater [18] is also glued to the crystal for the offline correction of thermal gain drift caused by temperature variation of the individual bolometer.

We put significant effort into the selection and handling of the detector materials with the objective of minimizing the background contamination for CUORE-0. In collaboration with the TeO₂ crystal grower at the Shanghai Institute of Ceramics, Chinese Academy of Sciences, we developed a radiopurity control protocol [19] to limit bulk and surface contaminations introduced in crystal production. Only materials certified for radiopurity were used to grow the crystals. After production, the crystals were transported to LNGS at sea level to minimize cosmogenic activation. Upon arrival at LNGS, a few crystals from each batch were instrumented as bolometers for characterization tests. For ²³⁸U (²³²Th) decay chain, the measured bulk and surface contaminations are less than 6.7×10^{-7} Bq/kg (8.4×10^{-7} Bq/kg) and $8.9 \times 10^{-9} \text{ Bq/cm}^2$ (2.0 × 10⁻⁹ Bq/cm²) at 90 % C.L., respectively [20]. Material screening of small parts, including NTD thermistors and silicon heaters, indicates that their radioactive content contributes to less than 10% of the total background in the 0vDBD region of interest (ROI).

Based on the experience of Cuoricino, we expect the most significant background contributions to come from the tower frame and the surrounding cylindrical thermal shield, both of which are made from radiopure electrolytic tough pitch copper [21]. Relative to Cuoricino, the total mass and surface area of the tower frame of CUORE-0 was reduced by a factor of 2.3 and 1.8, respectively. Monte Carlo studies predict a factor of 1.3 decrease in α background from the thermal shield arriving at the crystals due to the change in the geometry [22]. To further mitigate the surface contamination of the copper structure, we tested three surface treatment techniques [23] and chose a series of tumbling, electropolishing, chemical etching, and magnetron plasma etching for the surface treatment. The upper limit of the surface contamination of the cleaned copper was measured in R&D bolometers to be 1.3×10^{-7} Bq/cm² (90 % C.L.) for both ²³⁸U and ²³²Th [23].

The CUORE-0 detector assembly procedure was designed to minimize the recontamination of clean components. Tower assembly takes place in a dedicated class 1000 clean room in Hall A of the LNGS underground facility. To minimize exposure to radon (and radon progeny) in air, all steps of the assembly were performed under nitrogen atmosphere inside glove boxes [24]. All tools used inside the glove boxes, and especially those that would physically touch the detector components, were cleaned and certified for radiopurity. The assembled tower was enclosed in a copper thermal shield and mounted in the Cuoricino cryostat. To minimize exposure to the environment during mounting to the cryostat, mounting was performed in the Cuoricino clean room, and the tower was kept under nitrogen flux for as long as possible.

CUORE-0 uses for the first time flexible printed circuit board (PCB) cables and in situ wire bonding for electrical wiring of the tower. This is one of the major upgrades that significantly improved the robustness of bolometer readout wiring compared to the Cuoricino design. A set of flexible PCB cables with copper traces [25,26] was attached to the copper frame from the bottom plane to the top. The lower ends of the PCB copper traces were bonded to the metal contact pads of the thermistors and heaters using 25 μ m diameter gold wires. The upper ends of the PCB cables were connected through another custom-made flexible PCB at the 10 mK plate to a set of Manganin twisted pair flat ribbon cables running un-interrupted to the feedthroughs on the top plate of the cryostat. Overall, only 3 bolometers (6%) are not fully functional from the loss of 1 thermistor and 2 heaters. The two heater-less bolometers can be used in non-standard analysis without thermal gain correction in the future.

CUORE-0 is operated in the same cryostat, uses the same external lead and borated-polyethylene neutron shielding, and is enclosed in the same Faraday cage that was used for Cuoricino [12,27]. The front-end electronics [28–30] and data acquisition hardware are also identical to those used in Cuoricino. We implemented a new automated bias voltage scanning algorithm to locate the optimal working point that maximizes the signal-to-noise ratio (SNR). The bolometer signals are amplified and then filtered with six-pole Bessel low-pass filters. Subsequently, signals are digitized by two 32-channel National Instruments PXI analog-to-digital converters with a 125 S/s sampling rate, 18-bit resolution, and 21 V full scale. All samples are stored continuously on disk. Afterwards, in almost real-time, a constant fraction analysis trigger identifies triggered pulses with 626 sampling points (5.008 s), including a pre-trigger segment of 125 samples. Each bolometer has an independent trigger threshold, ranging from 50 to 100 keV. In addition to the signal triggers, each bolometer is pulsed periodically at 300 s intervals with a fixed and known energy through the heater. These "pulser" events are used to monitor and correct the gain of the bolometers [31]. Finally, a baseline trigger identifies a baseline pulse every 200 s to provide snapshots of the detector working temperatures and noise spectra.

CUORE-0 data are grouped into "data sets". Each data set consists of a set of initial calibration runs, a series of physics runs, and a set of final calibration runs. Calibration data refers to the sum of all calibration runs, while background data refers to the sum of all physics runs to search for 0ν DBD. During calibrations, the detector is irradiated using two thoriated tungsten wires, each with a ²³²Th activity of 50 Bq. The wires are inserted into two vertical tubes on

opposite sides of the tower that run between the outer vacuum chamber and the external lead shielding. We calibrate each channel using γ rays from daughter nuclei of ²³²Th in the energy range from 511 to 2615 keV. The signal rates on each bolometer for the calibration and background data are 60–70 and 0.5–1.0 mHz, respectively.

3 CUORE-0 performance and background

The CUORE-0 data reported in this article was collected between March and September 2013, with interruptions for dilution refrigerator maintenance. To account for temporary degraded performances on each individual bolometer due to large baseline excursions or elevated noise levels, we reject low-quality data intervals on a channel-by-channel basis. Consequently, the total exposure is obtained by summing the individual exposures of each bolometer. The accumulated TeO₂ exposure on 49 fully active channels is 7.1 kg · y for a ¹³⁰Te isotopic exposure of 2.0 kg · y, excluding all low-quality data intervals.

CUORE-0 data analysis follows the same procedure as was used for Cuoricino [12]. This includes amplitude evaluation, gain correction, energy calibration, and time coincidence analysis among the bolometers. Pulse amplitude is evaluated by first maximizing the SNR with an optimum filter. Fourier components of each pulse are weighted at each frequency by the expected SNR, which is calculated for each channel with an average pulse of 2615 keV γ rays and the average power spectra of the noise events. For the gain correction of each bolometer, the amplitudes of pulser events are fit against their baseline voltages to determine the gain dependence on temperature, and this temperature dependence is backed out for each signal pulse. For the energy calibration, we use a third order polynomial fit in the energy range from 0 to 3.9 MeV since the relationship between energy and stabilized amplitude is found to be slightly nonlinear. The deviation from a linear fit is less than 10 keV at the 2615 keV peak. If any two or more crystals register signal pulses within 100 ms of each other, the events are tagged as coincidence events. These are mostly attributed to backgrounds such as Compton-scattered γ rays or α decays on the surface of two adjacent crystals.

The event selection criteria can be categorized as follows: basic data quality, pile-up, pulse shape, and anti-coincidence. The basic data quality cut rejects events within low-quality data intervals, as mentioned in the beginning of the section. The pile-up cut requires that only one pulse exists in a 7.1 s window around the measured trigger time (see Fig. 1c). Due to the relatively long rise and decay times of a pulse and negligible pulse shape dependence on energy at energies above 1 MeV, the pulse shape of the possible ¹³⁰Te 0vDBD signal is expected to be similar to that obtained from the 2615 keV γ -ray peaks. Therefore, the pulse shape cut requires that the signal shape is comparable to that obtained from the average pulse recorded with 2615 keV γ -ray events, and that the pre-trigger baseline slope is smaller than 0.1 mV/Sample. The anti-coincidence cut requires that no other pulse in coincidence is recorded in the entire tower.

We evaluate the selection efficiency mainly using the 2615 keV ν -ray peak since it offers sufficient statistics at the energy closest to the ROI. However, since the 2615 keV γ -ray events occasionally occur in coincidence with other physical events, the efficiency of anti-coincidence cut was evaluated using the 1461 keV γ rays from ⁴⁰K decay, which are truly individual events. The selection efficiency was averaged over all active channels. The efficiency was obtained by first evaluating the slowly varying background rate under the peak by counting the number of events in the energy regions between 3 and 15σ above and below the peak. The background rate was then subtracted from the peak rate which was measured by counting the number of events within $\pm 3\sigma$ of the peak. The result was cross-checked by fitting the combined peak and background region $(\pm 15\sigma)$ with a Gaussian plus linear function. The difference between the two methods was integrated as the systematic uncertainty of the selection efficiency. The obtained efficiency is 92.9 ± 1.8 %, which is the efficiency of all cuts other than the anti-coincidence cut, obtained from the 2615 keV γ -ray peak, multiplied by the efficiency of the anti-coincidence cut, obtained from the 1461 keV peak after applying all other cuts, as described above. Since we are considering only single crystal events, we must include the confinement efficiency, i.e. the probability that both 0ν DBD electrons are contained inside the single crystal. This probability has been estimated by simulation to be $87.4 \pm 1.1\%$ [12]. Taking into account the 99.00 ± 0.01 % signal trigger efficiency, which is evaluated on pulser events, the total 0vDBD detection efficiency of CUORE-0 is 80.4 ± 1.9 %. This result is compatible with the value obtained from Cuoricino, which was found to be $82.8 \pm 1.1 \%$ [12].

The top panel in Fig. 2 shows the energy spectrum obtained using the ²³²Th calibration source. The spectrum is the sum of all 49 fully functional channels. The histogram in the bottom panel shows the background spectrum of CUORE-0 for the analysis presented. The presence of the pronounced γ -ray peaks from ²¹⁴Bi decay, a daughter nucleus of ²²²Rn, is attributed to the inclusion of data taken without the nitrogen gas purge in the Faraday cage around the dilution refrigerator. The nitrogen purge effectively suppresses ²¹⁴Bi γ -ray intensities by more than a factor of 5. The energy resolution in the ROI is defined as the FWHM of 2615 keV γ -ray peak, determined by a fit to the summed background spectrum of all fully functional channels. The median resolution of the individual channels in the calibration data is 6.0 keV, with a mean of 6.8 keV and a root mean

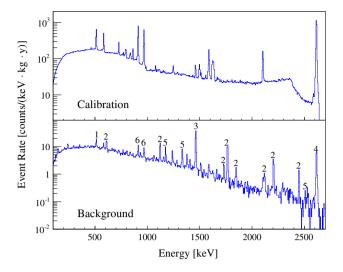


Fig. 2 CUORE-0 calibration (*top panel*) and background spectrum (*bottom panel*) over the data taking period presented in this work. γ -ray peaks from known radioactive sources in the background spectrum are labeled as follows: (1) e^+e^- annihilation; (2) ²¹⁴Bi; (3) ⁴⁰K; (4) ²⁰⁸Tl; (5) ⁶⁰Co; and (6) ²²⁸Ac

square deviation of 2.1 keV. The relatively larger mean and root mean square deviation are attributed to a few underperforming channels. The energy resolution of calibration data slightly deteriorates due to accidental coincidence events and a higher overall noise level compared to background data.

When compared to Cuoricino, one new noise contribution is correlated microphonic noise on multiple channels introduced by the new flexible PCB wiring. The vibration of one PCB cable might introduce common-mode noise in all the channels on that cable, which is apparent in the lowfrequency part of the signal band and degrades the energy resolution. However, even with the correlated noise in play, the energy resolution of 5.7 keV (FWHM) is better, on average, than that of Cuoricino. Furthermore, we have on-going studies seeking to improve the energy resolution by regressing the correlated noise out of the bolometer signals [32].

As of this writing, we have kept the 0 ν DBD region blinded. Our blinding procedure is a form of *data salting*, where we randomly exchange a blinded fraction of events within ± 10 keV of the 2615 keV γ -ray peak with events within ± 10 keV of the 0 ν DBD Q-value. The exchange probability varies between 1 and 3 % and is randomized run by run. Since the number of 2615 keV γ -ray events is much larger than that of possible 0 ν DBD events, the blinding algorithm produces an artificial peak around the 0 ν DBD Q-value and blinds the real 0 ν DBD rate of 130 Te. This method of blinding the data preserves the integrity of the possible 0 ν DBD events while maintaining the spectral characteristics with measured energy resolution and introducing no discontinuities in the spectrum.

The background rate in the ROI is evaluated using the blinded spectrum in the energy range 2470–2570 keV. This

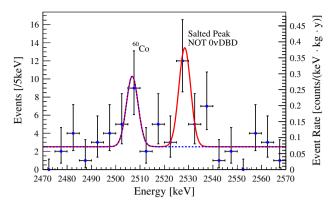


Fig. 3 Blinded energy spectrum and 0vDBD global fit in the region of interest. The unbinned maximum-likelihood fit is shown in solid *red line*. The *dotted blue line* illustrates the peak from ⁶⁰Co and the linear background only to highlight the difference between salted 0vDBD peak and background. The flat background from the fit is $0.071 \pm 0.011 \text{ counts}/(\text{keV} \cdot \text{kg} \cdot \text{y})$. ⁶⁰Co peak position, salted peak position, and rate are $2506.8 \pm 1.1 \text{ keV}$, $2528.4 \pm 1.0 \text{ keV}$, and $1.3 \pm 0.5 \text{ counts}/(\text{keV} \cdot \text{kg} \cdot \text{y})$, respectively

region includes the ⁶⁰Co sum-peak at 2506 keV and the salted peak at the 0vDBD O-value, as shown in Fig. 3. We use an unbinned maximum-likelihood fit to estimate the background rate in the ROI. The likelihood function consists of the sum of a ⁶⁰Co Gaussian peak, a salted 0vDBD Gaussian peak, and a flat background. In the fit, the mean of the 60 Co peak is initialized to 2506 keV and the mean of the salted 0vDBD peak at 2528 keV. The FWHM of both peaks is fixed to the detector resolution at 5.7 keV. As shown in Fig. 3, the fit reveals that the overall background rate in the ROI is 0.071 ± 0.011 (stat) counts/(keV \cdot kg \cdot y). For comparison, the background rate of the Cuoricino crystals with the same dimension is 0.153 ± 0.006 counts/(keV · kg · y). Systematic uncertainties arising from background shape are studied by comparing a constant and a linear background models, and are found to be less than 3%. The systematic contribution from the uncertainty in energy calibration is less than 1%.

The two major sources of background in the ROI are degraded α particles from surface contamination on the detector components and γ rays that originate from the cryostat. Degraded α particles with a decay energy of 4 to 8 MeV may deposit part of their energy in the 0vDBD ROI. These α events form a continuous energy spectrum extending from their decay energy to well below 0vDBD region. The α background rate in the ROI is estimated by counting events in the " α flat continuum region", which is defined to be from 2.7 to 3.9 MeV (excluding the ¹⁹⁰Pt peak region from 3.1 to 3.4 MeV). This energy range is above almost all naturally occurring γ rays, in particular the 2615 keV γ rays from ²⁰⁸Tl decay. Figure 4 shows the background energy spectrum of CUORE-0 (shaded red) and Cuoricino (black). The measured rate for CUORE-0 is

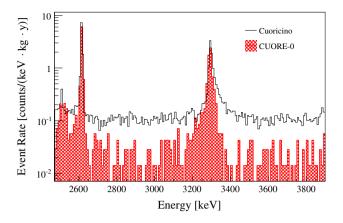


Fig. 4 Background spectrum of CUORE-0 (*red* with shades) and Cuoricino (*black*) in the region dominated by degraded α particles. The figure shows reduction of the flat background caused by degraded α particles in the energy region of [2.7–3.1] and [3.4 – 3.9] MeV

 0.019 ± 0.002 counts/(keV \cdot kg \cdot y), which improves on the Cuoricino result (0.110 \pm 0.001 counts/(keV \cdot kg \cdot y)) by a factor of 6.

The γ -ray background in the ROI is predominantly Compton-scattered 2615 keV γ rays originating from ²⁰⁸Tl in the cryostat. Since CUORE-0 is hosted in the same cryostat as was used for Cuoricino, the γ -ray background is expected to be similar. The γ -ray background is estimated as the difference between overall background in the ROI and the degraded α background in the continuum. The measured γ -ray backgrounds of CUORE-0 and Cuoricino are indeed compatible [12], consistent with the hypothesis that the background in the ROI is composed of γ rays from the cryostat and degraded α particles.

4 Projected sensitivity of CUORE-0

Using the measured background rate and energy resolution of the 2615 keV γ -ray peak, we obtain the CUORE-0 sensititvity with the approach outlined in [33]. With the excellent energy resolution, we construct a single-bin counting experiment with a 5.7 keV bin centered at the 0 ν DBD Q-value. The sensitivity is obtained by comparing the expected number of signal events with Poissonian fluctuations from the expected background rate in this bin. Figure 5 shows the 90% C.L. sensitivity of CUORE-0. With one year of live time, or 11 kg · y isotope exposure, CUORE-0 is expected to surpass the ¹³⁰Te 0 ν DBD half-life sensitivity achieved by Cuoricino, 2.8 × 10²⁴ y.

5 Summary and outlook

We present the energy resolution and background measurements of CUORE-0 detector from the 7.1 kg \cdot y exposure accumulated up to September 2013. The measured 5.7 keV FWHM in the 0 ν DBD ROI represents a slight improvement

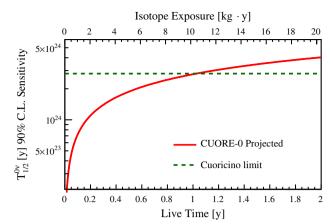


Fig. 5 Sensitivity of CUORE-0 with the measured background rate in the ROI of $0.071 \text{ counts}/(\text{keV} \cdot \text{kg} \cdot \text{y})$ and energy resolution of 5.7 keV FWHM. The CUORE-0 sensitivity is expected to surpass that of Cuoricino with 1 year of live time

over Cuoricino and validates the CUORE-0 wiring scheme and assembly procedure. The background rates have been measured to be $0.071 \pm 0.011 \text{ counts}/(\text{keV} \cdot \text{kg} \cdot \text{y})$ in the ROI and $0.019 \pm 0.002 \text{ counts}/(\text{keV} \cdot \text{kg} \cdot \text{y})$ in the α continuum region. These results are a factor of 2 and 6 improvement compared to Cuoricino, due to more rigorous copper surface treatment, improved crystal production and treatment protocols, and more stringent assembly procedures in the clean environment. The CUORE-0 sensitivity is expected to surpass that of Cuoricino with one year of live time.

As a technical prototype for CUORE, CUORE-0 demonstrates the feasibility of instrumenting an ultra-pure ton-scale bolometer array with 988 channels. By enhancing the procedure of the on-going CUORE assembly, we have improved assembly success rate to close to 100%, which is a crucial achievement for large arrays such as CUORE. We have started implementing the noise decorrelation algorithms into the CUORE-0/CUORE data analysis package, with the aim of further improving energy resolution. CUORE-0 reconfirms the effectiveness of the copper cleaning technique and clean assembly procedure developed for CUORE. Compared to CUORE-0, the larger array of CUORE affords more powerful time coincidence analysis and more effective selfshielding from external backgrounds, particularly those originating from the copper thermal shields or cryostat. With this stronger background rejection and the already demonstrated reduction of surface contamination, the CUORE background goal of 0.01 counts/(keV \cdot kg \cdot y) is expected to be within reach. The projected half-life sensitivity to 130 Te 0ν DBD is 9.5×10^{25} y (90 % C.L.) with 5 years of live time [33], reaching an effective Majorana neutrino mass sensitivity of 0.05 to 0.13 eV [5–11,34].

Acknowledgments The CUORE Collaboration thanks the directors and staff of the Laboratori Nazionali del Gran Sasso and the techni-

cal staff of our laboratories. This work was supported by the Istituto Nazionale di Fisica Nucleare (INFN); the Director, Office of Science, of the U.S. Department of Energy under Contract Nos. DE-AC02-05CH11231 and DE-AC52-07NA27344; the DOE Office of Nuclear Physics under Contract Nos. DE-FG02-08ER41551 and DEFG03-00ER41138; the National Science Foundation under Grant Nos. NSF-PHY-0605119, NSF-PHY-0500337, NSF-PHY-0855314, NSF-PHY-0902171, and NSF-PHY-0969852; the Alfred P. Sloan Foundation; the University of Wisconsin Foundation; and Yale University. This research used resources of the National Energy Research Scientific Computing Center (NERSC).

Open Access This article is distributed under the terms of the Creative Commons Attribution License which permits any use, distribution, and reproduction in any medium, provided the original author(s) and the source are credited.

Funded by SCOAP³ / License Version CC BY 4.0.

References

- I. Avignone, T. Frank, S.R. Elliott, J. Engel, Rev. Mod. Phys. 80, 481 (2008). arXiv:0708.1033
- M. Auger et al. (EXO Collaboration), Phys. Rev. Lett. 109, 032505 (2012), arXiv:1205.5608.
- A. Gando et al. (KamLAND-Zen Collaboration), Phys. Rev. Lett. 110, 062502 (2013). arXiv:1211.3863.
- M. Agostini et al. (GERDA Collaboration), Phys. Rev. Lett. 11, 122503 (2013). arXiv:1307.4720
- J. Menendez, A. Poves, E. Caurier, F. Nowacki, Nucl. Phys. A 818, 139 (2009). arXiv:0801.3760
- 6. P.K. Rath et al., Phys. Rev. C 87, 014301 (2013)
- T.R. Rodriguez, G. Martinez-Pinedo, Phys. Rev. Lett. 105, 252503 (2010). arXiv:1008.5260
- D.-L. Fang, A. Faessler, V. Rodin, F. Simkovic, Phys. Rev. C 83, 034320 (2011). arXiv:1101.2149
- A. Faessler, V. Rodin, F. Simkovic, J. Phys. G: Nucl. Part. Phys. 39, 124006 (2012). arXiv:1206.0464
- J. Suhonen, O. Civitarese, J. Phys. G: Nucl. Part. Phys. 39, 124005 (2012)

- J. Barea, J. Kotila, F. Iachello, Phys. Rev. C 87, 014315 (2013). arXiv:1301.4203
- E. Andreotti et al. (Cuoricino Collaboration), Astropart. Phys. 34, 822 (2011). arXiv:1012.3266
- M. Redshaw, B.J. Mount, E.G. Myers, F.T. Avignone III, Phys. Rev. Lett. 102, 212502 (2009). arXiv:0902.2139
- 14. N.D. Scielzo et al., Phys. Rev. C 80, 025501 (2009). arXiv:0902.2376
- 15. S. Rahaman et al., Phys. Lett. B 703, 412 (2011)
- M.A. Fehr, M. Rehkamper, A.N. Halliday, Int. J. Mass Spectrom. 232, 83 (2004)
- 17. R. Ardito et al. (CUORE Collaboration) (2005). arXiv:hep-ex/0501010
- 18. A. Alessandrello et al., Nucl. Instrum. Meth. A 412, 454 (1998)
- C. Arnaboldi et al., J. Cryst. Growth 312, 2999 (2010). arXiv:1005.3686
- F. Alessandria et al. (CUORE Collaboration), Astropart. Phys. 35, 839 (2012). arXiv:1108.4757
- 21. http://www.aurubis.com
- 22. D.R. Artusa et al. (CUORE Collaboration), in preparation
- F. Alessandria et al. (CUORE Collaboration), Astropart. Phys. 45, 13 (2013). arXiv:1210.1107
- 24. M. Clemenza, C. Maiano, L. Pattavina, E. Previtali, Eur. Phys. J. C 71, 1 (2011)
- 25. C. Brofferio et al., Nucl. Instrum. Meth. A 718, 211 (2013)
- 26. E. Andreotti et al., JINST 4, P09003 (2009)
- C. Arnaboldi et al. (Cuoricino Collaboration), Phys. Rev. C 78, 035502 (2008). arXiv:0802.3439
- 28. C. Arnaboldi et al., Nucl. Instrum. Meth. A 617, 327 (2010)
- 29. C. Arnaboldi et al., Nucl. Instrum. Meth. A 520, 578 (2004)
- 30. C. Arnaboldi et al., IEEE Trans. Nucl. Sci. 49, 2440 (2002)
- C. Arnaboldi, G. Pessina, E. Previtali, IEEE Trans. Nucl. Sci. 50, 979 (2003)
- 32. C. Mancini-Terracciano, M. Vignati, JINST 7, P06013 (2012). arXiv:1203.1782
- 33. F. Alessandria et al. (CUORE Collaboration) (2011). arXiv:1109.0494
- 34. J. Kotila, F. Iachello, Phys. Rev. C 85, 034316 (2012). arXiv:1209.5722